

# Classical Mechanics (20%)

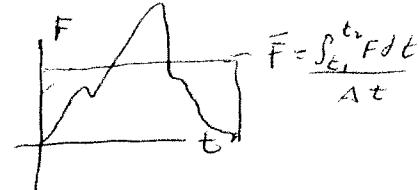
Dan Elton

$$W = \int_{x_1}^{x_2} F(x) dx$$

$$\vec{F} = -\vec{\nabla} V$$

## Impulse

$$\vec{J} = \int_{t_1}^{t_2} \vec{F} dt = \vec{F} \Delta t = \Delta \vec{P}$$



## Center of mass

$$CM = \frac{1}{M} \sum_{i=1}^n r_i m_i$$

## Air resistance

$$\text{Low velocity: } \vec{F}_r = -bv$$

$$\text{high velocity: } \vec{F}_r = -cv^2$$

## Rocket motion

$$F_{\text{thrust}} = \frac{dp}{dt} = V_{\text{exhaust}} \frac{dm}{dt}$$

## Collisions

### Elastic collisions

$$V_{1f} = \left( \frac{m_1 - m_2}{m_1 + m_2} \right) V_{1i} + \left( \frac{2m_2}{m_1 + m_2} \right) V_{2i}$$

$$V_{2f} = \left( \frac{2m_1}{m_1 + m_2} \right) V_{1i} + \left( \frac{m_2 - m_1}{m_1 + m_2} \right) V_{2i}$$

notes: in the case  $m_1 = m_2$ ,

$$V_{1f} = V_{2i}$$

~~V<sub>2f</sub>~~ -

in the case  $m_2 \gg m_1$ ,  $V_{2i} = 0$

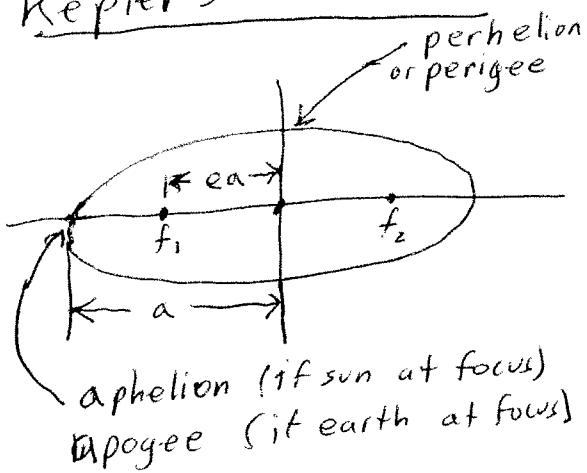
$$V_{2f} \approx -V_{1i}$$

In elastic collisions, the relative velocities of the particles is equal and opposite to the relative velocities afterwards. ( $\Delta V_{rel} = V_{2i} - V_{1i}$  is conserved)

## Gravitation

$$\boxed{\vec{F} = G \frac{m_1 m_2}{r^2} \hat{r}}$$

## Kepler's 1st Law



## eccentricity

$e \leq 0$	circle
$0 < e < 1$	ellipse
$e = 1$	parabola
$e > 1$	hyperbola

## Kepler's 2nd Law

$$\frac{dA}{dt} = \lim_{\Delta t \rightarrow 0} \frac{\Delta A}{\Delta t} = \lim_{\Delta t \rightarrow 0} \frac{\frac{1}{2} r \cdot r \Delta \theta}{\Delta t} = \frac{1}{2} r^2 \omega$$

$$\boxed{L = mvr = mr^2\omega = \text{const.}}$$

## Kepler's 3rd Law

$$T^2 = \left(\frac{4\pi^2}{GM}\right) r^3$$

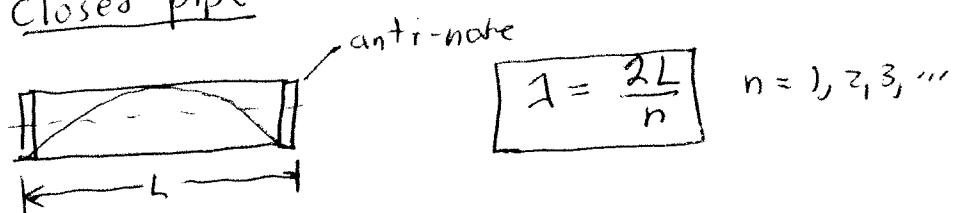
$$\boxed{T^2 \propto r^3}$$

## Sound Waves

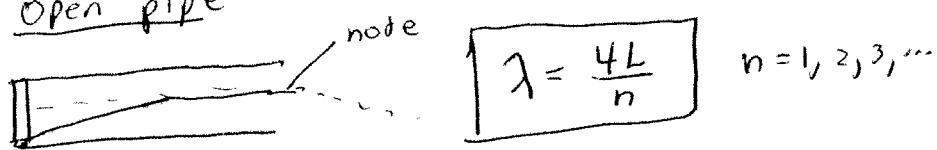
$$V_{\text{sound, air}} = 343 \text{ m/s} = 1 \text{ mile / 5 seconds}$$

$$V_{\text{sound, air}} \propto \sqrt{T'}$$

### Closed pipe



### Open pipe



## Doppler effect

$v_o$  = source frequency

$v'$  = observer frequency

$v$  = speed of sound

$v_s$  = speed of source

$v_o$  = speed of observer

master

$$v' = v_o \frac{v \pm v_o}{v + v_s}$$

observer  
at rest

$$v' = v_o \frac{v}{v + v_s}$$

# Angular Momentum

$$\boxed{L = m\vec{v} \times \vec{r} = \vec{p} \times \vec{r} = \vec{r} \times \vec{p} = mrV_i = I\omega}$$

## Rotational Analogs

$$\boxed{K_{\text{rot}} = \frac{1}{2}I\omega^2}$$

$$\boxed{\gamma = I\alpha}$$

## Moments of Inertia



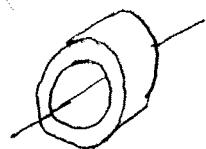
$$\boxed{I = mR^2}$$

hoop



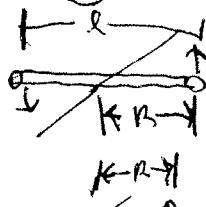
$$\boxed{I = \frac{mR^2}{2}}$$

disk / cylinder

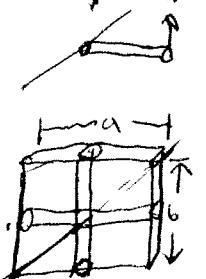


$$\boxed{I = \frac{m(R_1^2 + R_2^2)}{2}}$$

hollow disk / cylinder



$$\boxed{I = \frac{mL^2}{12}}$$

+ rod through center of mass \*  $\left( = \frac{mR^2}{3} \right)$ 

$$\boxed{I = \frac{mL^2}{3}}$$

\* rod through one end \*

$$\boxed{I = \frac{m(a^2 + b^2)}{12}}$$

solid plate



$$\boxed{I = \frac{2}{5}mR^2}$$

sphere

## \* Parallel Axis Theorem \*

$$\boxed{I = I_{cm} + md^2}$$

# Classical Mechanics: Oscillations

## SHO

$$\ddot{x} + \frac{k}{m}x = 0$$

$$\omega^2 = k/m$$

$$\Rightarrow f = \frac{1}{2\pi} \sqrt{\frac{k}{m}}$$

## Simple pendulum

$$\tau = l m g \sin\theta = m l^2 \dot{\theta}' = I \dot{\theta}'$$

$$\Rightarrow \ddot{\theta}' + \frac{g}{l} \sin\theta' = 0 \quad \text{or} \quad \ddot{\theta}' + \frac{l m g \sin\theta'}{I} = 0$$



$$\Rightarrow T = 2\pi \sqrt{\frac{l}{g}}$$

## Physical pendulum

$$T = 2\pi \sqrt{\frac{I}{mgL}}$$

## Forced/damped Oscillations

$$m\ddot{x} + b\dot{x} + kx = 0$$

$$\ddot{x} + 2\beta\dot{x} + \omega_0^2 x = f_0 \cos\omega t$$

$$\Rightarrow x(t) = A \cos(\omega t - \delta)$$

$$A = \sqrt{\frac{f_0^2}{(\omega_0^2 - \omega^2)^2 + 4\beta^2\omega^2}}$$

$$\delta = \arctan \left( \frac{2\beta\omega}{\omega_0^2 - \omega^2} \right)$$

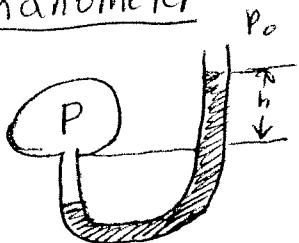
## Fluid statics

$$P = \frac{dF}{dA}$$

Pascal's principle : basis for hydraulics

Archimedes's principle : buoyant force = weight of displaced water

## manometer



$$P - P_0 = \rho g h$$

## Fluid Dynamics

### Continuity Equation

$$P \vec{\nabla} \cdot \vec{V} + \frac{dp}{dt} = 0$$

### Bernoulli's Equation

$$P + \frac{1}{2} \rho v^2 + \rho gy = \text{constant}$$

# Intermediate Mechanics

$$L = T - U$$

$$H = T + V$$

$$L - \sum_j \dot{q}_j \frac{\partial L}{\partial \dot{q}_j} = H - \sum_j p_j \dot{q}$$

Lagrange-Euler Equation (w/ Lagrange multipliers)

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} + \sum_k \lambda_k \frac{\partial f}{\partial q} = 0$$

## Hamilton's Equations

$$\dot{q} = \frac{\partial H}{\partial p}$$

$$\dot{p} = -\frac{\partial H}{\partial q}$$

# E & M (18%)

converted

## The Electric field

$$\vec{E} = \frac{\vec{F}}{q_0}$$

$$\vec{D} = \epsilon_0 \vec{E} = \vec{P}$$

## Electric susceptibility / Dielectric

$$E \propto \frac{1}{\epsilon_0} \quad \underline{\epsilon = K\epsilon_0}$$

$K$  = dielectric constant

## Electric Field of a Dipole

$$E = \frac{1}{4\pi\epsilon_0} \frac{qd}{[x^2 + (\frac{d}{2})^2]^{3/2}} \approx \frac{1}{4\pi\epsilon_0} \frac{qd}{x^3} \text{ when } x \gg d$$

## Dipole moment

$$\vec{p} = q\vec{d}$$

## Electric dipole torque

$$\vec{\tau} = \vec{p} \times \vec{E}$$

## Electric Field of an Infinite Sheet

$$\vec{E}_z = \frac{\sigma}{2\epsilon_0}$$

## Electric Potential Energy

$$U(r) = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r}$$

## Electric Potential

$$V = \frac{U}{q_0}$$

$$\Delta V = \int_a^b \vec{E} \cdot d\vec{s} \Leftrightarrow \vec{E} = -\vec{\nabla}V$$

## Capacitance

$$C = \frac{Q}{V}$$

## Parallel Plate Capacitor

$$C = \frac{\epsilon_0 A}{D}$$

$$V = \int_a^b E \cdot ds = \int \frac{q}{A \epsilon_0}$$

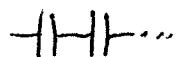
$$E = \frac{\sigma}{2\epsilon} + \frac{\sigma}{2\epsilon} = \frac{\sigma}{\epsilon_0} = \frac{Q}{A \epsilon_0}$$

## Capacitors in Parallel



$$C_{\text{parallel}} = C_1 + C_2 + \dots$$

## Capacitors in Series



$$\frac{1}{C_{\text{series}}} = \frac{1}{C_1} + \frac{1}{C_2} + \dots$$

## Energy stored in a Capacitor

$$U = \frac{1}{2} CV^2 = \frac{1}{2} \frac{Q^2}{C}$$

## Capacitor w/ dielectric

$$C = \kappa C_0$$

## OHM's Law

$$V = IR \quad I = \frac{V}{R} \quad R = \frac{V}{I}$$

## Resistivity

$$\rho = \frac{V}{J} = \frac{V}{I/A}$$

$$\frac{V}{J} = \frac{IR}{J}$$

$$\rho = R \frac{A}{l}$$

$$R = \rho \frac{l}{A}$$

## Inductance

$$E = L \frac{di}{dt}$$

Faraday's Law

$$L = \frac{\Phi}{i}$$

## Energy of an inductor

$$U = \frac{1}{2} L i^2$$

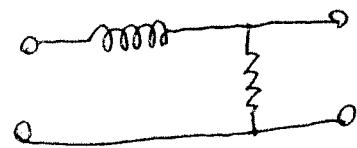
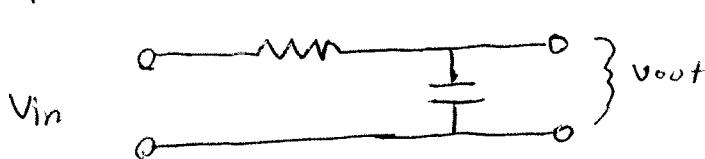
$$V_{out} = IR = \frac{V_{in}}{Z} R = \frac{VR}{R + i\omega L}$$

as  $\omega \rightarrow \infty$ ,  $V_{out} \rightarrow 0$   
 (break low pass filter)

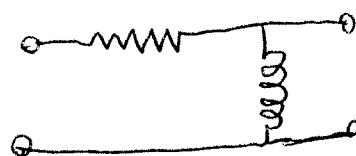
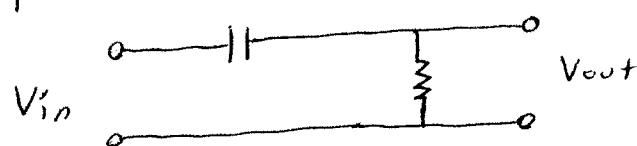
~~$$V_{out} = \frac{V_{in} R}{R + i\omega L} = V_{in} R / R$$~~

## Electric Filters

### low pass filters



### high pass filters



## Current Density

$$j = \frac{i}{A} = V_s n e$$

Impedance:  $V = I Z$

defined as  $V/I$ ; as a function of  $\omega$ . It is analogous to resistance with DC.

### Reactance:

Is the complex part of impedance and results in a phase shift of the AC signal.

Resistive.

## Hall effect coefficient

$$R_H = -\frac{1}{nec} = \frac{E_y}{j_x B}$$

## Impedance / AC circuit Analysis

$$Z_{inductor} = j\omega L$$

$$Z_{capacitor} = \frac{-j}{\omega C}$$

$$Z_{resistor} = R$$

see  
next  
page

~~Master equation~~:  $Z = R + j\omega L - j/\omega C$

$$V = I Z$$

## Electric Potentials

$$\vec{E} = -\nabla V - \frac{\partial \vec{A}}{\partial t}$$

$$\vec{B} = \nabla \times \vec{A}$$

## Energy of E & B Fields

$$U = \frac{1}{2} \int_V (\epsilon_0 E^2 + \frac{1}{\mu_0} B^2) dV$$

## Momentum of E & B Fields

$$\vec{P} = \epsilon_0 \int_V \vec{E} \times \vec{B} dV$$

## Poynting Vector

$$\vec{S} = \frac{1}{\mu_0} (\vec{E} \times \vec{B})$$

## Biot - Savart Law

$$\vec{B} = \frac{\mu_0}{4\pi} \int I \frac{d\vec{l} \times \vec{dr}}{r^2} = \frac{\mu_0}{4\pi} \int I \frac{d\vec{x} \times \vec{r}}{r^3}$$

## Lamor Formula

$$P = \frac{q^2 \alpha^2}{8\pi r_0^2}$$

$$P \propto q^2 \alpha^2$$

## Maxwell's Equations

### Gauss's Law

$$\oint_{\partial V} \vec{E} \cdot d\vec{A} = \frac{\rho_{\text{enc}}}{\epsilon_0}$$

$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

### No Magnetic monopoles

$$\oint_{\partial V} \vec{B} \cdot d\vec{A} = 0$$

$$\nabla \cdot \vec{B} = 0$$

### Faraday's Law

$$\oint \vec{E} \cdot d\vec{L} = - \frac{\partial \Phi_E}{\partial t}$$

$$\nabla \times \vec{E} = - \frac{\partial \vec{B}}{\partial t}$$

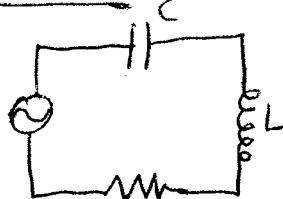
### Ampere's Law

$$\oint \vec{B} \cdot d\vec{L} = \mu_0 i_{\text{enc}} + \mu_0 \epsilon \frac{\partial \Phi_E}{\partial t}$$

$$\nabla \times \vec{B} = \mu_0 \vec{J} + \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t}$$

### AC Circuit Analysis (continued)

#### LRC circuit



Resonance occurs when the reactance (complex impedance) is 0.  
 i.e.,  $Z = R + j(\omega L - 1/\omega C) \Rightarrow \omega L = 1/\omega C \Rightarrow \omega^2 = \frac{1}{LC}$

#### Mechanical Analog

$$m\ddot{x} + b\dot{x} + kx = 0$$

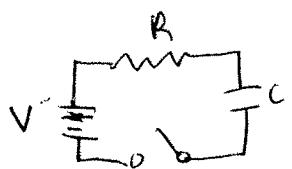
$$L\ddot{q} + R\dot{q} + \frac{q}{C} = 0$$

$$\omega = \sqrt{\frac{1}{LC}} = \frac{1}{\sqrt{LC}}$$

## AC Circuit Analysis, continued

$$V = IR$$

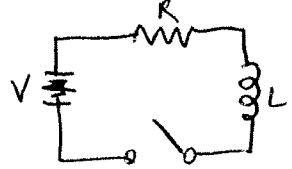
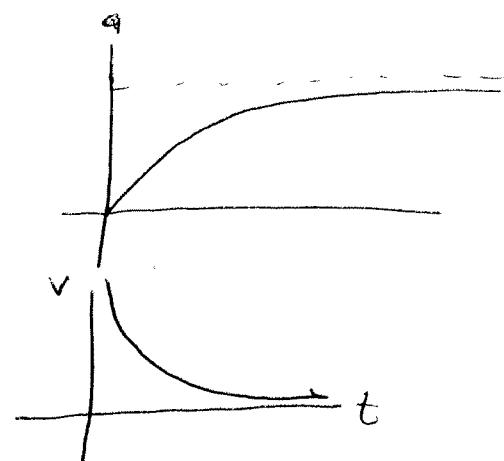
$$\boxed{\text{Power} = I^2 R}$$



$$\dot{q}R + \frac{q}{C} = V$$

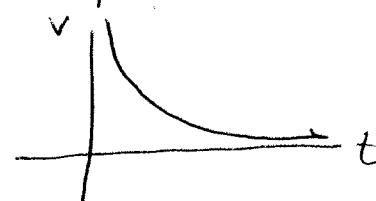
$$q = CV[1 - e^{-t/RC}]$$

$$C = \frac{a}{V}$$



$$\dot{q}'L + \dot{q}R = V$$

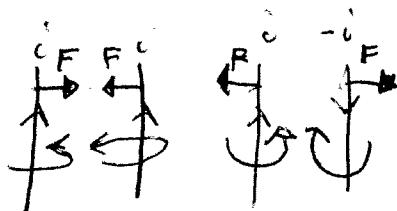
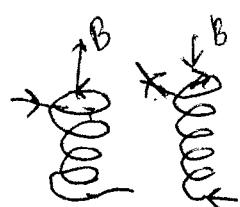
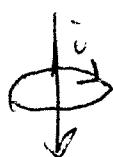
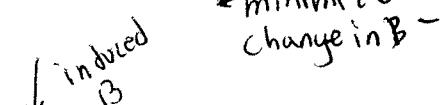
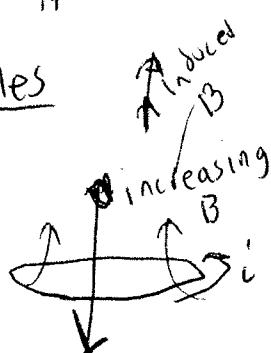
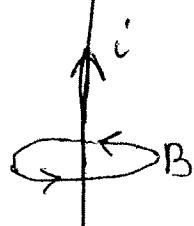
$$\dot{q}' + \frac{\dot{q}R}{L} = V$$



## Lenz's Law

"An induced current is always such to oppose the changes causing it"

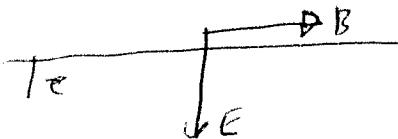
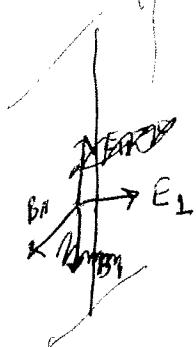
### Right-Hand Rule



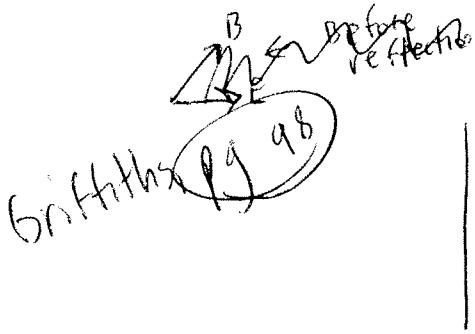
## More misc. E & M

Electric boundary conditions for reflection  
from an infinitely long, perfectly conducting sheet

$$\vec{E}_{||} = 0, \quad \vec{B}_{\perp} = 0$$



for a perfect conductor, there  
can never be an  $E$ -field parallel  
to the surface. The phase of an incoming  
wave shifts by  $\pi$ , so the  $E$ -field  
reverses direction, canceling out.



### Cyclotron frequency

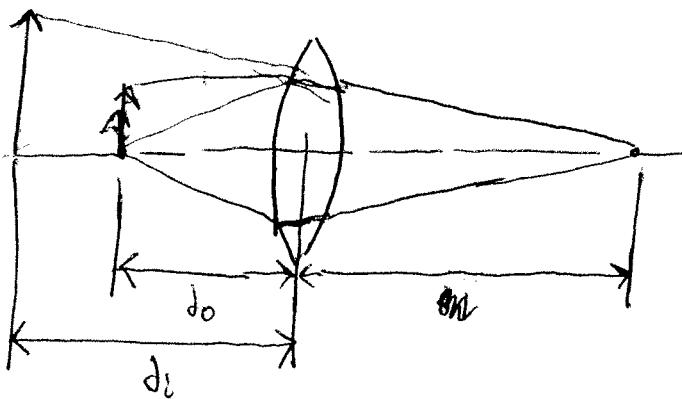
$$qVB = \frac{mv^2}{r}$$

$$v = \frac{qVB}{m} = \frac{2\pi r}{T} f \Rightarrow \omega = \frac{qB}{m}$$



# Optics & Wave Phenomena (9/.)

## Thin Lens Equation



$$\frac{1}{f} = \frac{1}{d_o} + \frac{1}{d_i}$$

$d_o$  = object distance

$d_i$  = image distance

$$M = -\frac{d_i}{d_o}$$

$$M = \frac{f_o}{f_e}$$

## Rayleigh Criterion

$$\theta = \frac{1.22\lambda}{D}$$

## Malthus's law for Polarizers

$$I(\theta) = I(0) \cos^2 \theta$$

## Light Intensity

$$I = \epsilon_0 c \langle E^2 \rangle_t$$

## Snell's Law

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

## Total internal reflection

$$\theta = \arcsin \left( \frac{n_2}{n_1} \right)$$

## Brewster's angle

$$\theta_B = \arctan\left(\frac{\theta_2}{\theta_1}\right)$$

## Optical Path Length

$$OPL = \int n dl = \sum_i n_i di$$

## Maxwell's Relation

$$n \approx \sqrt{K}$$

## Double slit interference /

$$ds \sin \theta = m\lambda$$

$m = 0, 1, 2, \dots$  (maxima)

$$s = \theta r \quad \theta \approx \frac{y}{r} \approx \tan \theta$$

$$\approx \sin \theta \approx \theta \approx \frac{y}{D}$$

for maxima

$$y \approx \frac{m\lambda D}{d}$$

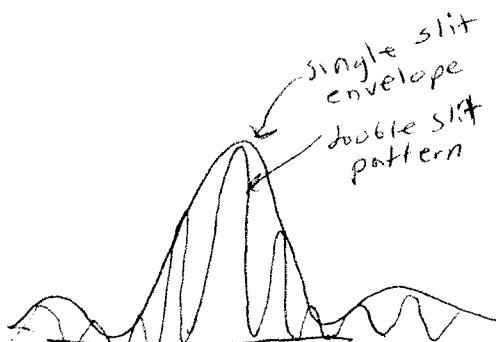
Phase change after reflectance  $\approx \pi$  (from lower to higher  $n$ )

## Bragg Diffraction

Waves are in phase when

$$2ds \sin \theta = n\lambda$$

$n = 0, 1, 2, \dots$



## Other diffraction problems

1-slit diffraction

minima

width at 1st

$$\sin \theta = m\lambda$$

Diffraction  
grating as well

double-slit diffraction

minima  $\sin \theta = m\lambda$

Circular diffraction, first minima

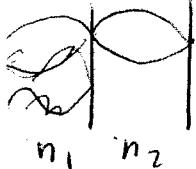
$$\sin \theta \approx \frac{1.22\lambda}{D}$$

(same as Rayleigh criterion)

"missing fringes" - occurs when diffraction minima cancel interference maxima in the double slit experiment

## Thin films

general principles:  
phase change occurs  
when  $n_2 > n_1$ , does  
not occur when  $n_2 \leq n_1$



constructive interference  $2t = \frac{\lambda}{2}$

destructive interference  $2t = \lambda$

# Thermodynamics : 10%

## Heat Capacity

$$C = \frac{\partial Q}{\partial T}$$

## Specific Heat Capacity

$$c = \frac{1}{m} \frac{\partial Q}{\partial T}$$

## Mayer's Equation

$$C_p - C_v = R = Nk$$

## Ideal Gas Law

$$PV = nRT$$

$$PV = NkT$$

## Work

$$W = - \int P dV$$

general equation

constant volume :  $W = 0$

constant pressure :  $W = -P(V_f - V_i)$

adiabatic :  $PV^\gamma = \text{const.} \rightarrow W = \frac{1}{\gamma-1} (P_f V_f - P_i V_i) \quad \gamma = \frac{C_p}{C_v}$

constant temperature  $W = -nRT \ln\left(\frac{V_f}{V_i}\right)$

## Thermal Expansion

$$\Delta L = \alpha L \Delta T$$

## Speed of sound

$$V_s \propto T^{1/2}$$

## Entropy

$$ds = \frac{\partial Q}{T}$$

## Boltzmann's Law

$$S = k \log w$$

w = # of possibilities

## Law of Dulong & Petit

$$C_v = 3R = 3Nk \quad \text{at high } T$$

## Debye $T^3$ Law

$$C_v \propto T^3 \quad \text{at lower } T$$

## \*Equipartition Theorem\*

$$E = \frac{1}{2} n k T$$

n = DOF  
for a solid,  
n = 6       $\begin{matrix} \rightarrow 3kT \\ \rightarrow 3PE \end{matrix}$

## specific heat of a metal

$$c = aT + B T^3$$

$$c = e^{aT} \quad (\text{superconducting})$$

## Stirling's Theorem/approximation

$$\ln N! \approx N \ln N - N \quad N \gg 0$$

## Thermodynamics pg 2

### Internal energy of a gas (partition theorem)

$$U = \left(\frac{1}{2} n k T\right) N$$

$n$  = degrees of freedom  
 $N$  = number of atoms/molecules

### Van der Waals equation

$$\left(p + a \frac{n^2}{V^2}\right)(V - nb) = nRT$$

### Maxwellian speed distribution

$$n(v) = 4\pi N \left(\frac{m}{2\pi kT}\right)^{3/2} v^2 e^{-mv^2/2kT}$$

### Maxwell-Boltzmann energy distribution

$$n(E) = \frac{2N}{\pi} \frac{1}{(kT)^{3/2}} E^{1/2} e^{-E/kT}$$

### Bose-Einstein distribution

$$f_{BE}(E) = \frac{1}{e^{(E-E_0)/kT} - 1}$$

indistinguishable  
no Pauli Exclusion

### Fermi-Dirac distribution

$$f_{FD}(E) = \frac{1}{e^{(E-E_0)/kT} + 1}$$

indistinguishable  
Pauli Exclusion principle

### Boltzmann Distribution / Maxwell-Boltzmann

$$f_j = \frac{N e^{-E_j/kT}}{\sum_i g_i e^{-E_i/kT}}$$

distinguishable

## Blackbody Planck Formula

$$\frac{\text{Power}}{\text{Area}, \lambda} = \frac{\text{Flux}}{\lambda} = \frac{dR}{d\lambda} = \frac{2\pi h c^2}{\lambda^5 (e^{h\lambda/kT} - 1)}$$

## Stephan-Boltzmann Law

$$R = \text{Power Output} = \sigma T^4$$

## Wein's Law for Blackbody

$$\lambda_{\max} \propto \frac{1}{T}$$

## First Law of Thermodynamics

A system goes from state i to f through various paths, the quantity  $Q + W$  is always the same.

$$dE = dQ + dW$$

## Second Law of Thermodynamics

There are no perfect heat engines.

→ It is not possible for a cyclical process to convert heat entirely to work.

→ It is impossible to build a heat engine more efficient

than a Carnot engine.  $\eta_{\text{Carnot}} = 1 - \frac{T_L}{T_H} = \frac{T_H - T_L}{T_H}$

→ A perfectly reversible engine has Carnot efficiency.

→ It is impossible to reach absolute zero. (Sometimes called "3rd law")

## Zeroth Law of Thermodynamics

If A & B are in T, E, then with C<sub>j</sub>

then A and B are in T, E, with each others

## Fourier's Law of Heat Conduction

$$\vec{q} = -k \nabla T$$

$k$  = thermal conductivity  
 $q$  = heat flux,  $= \frac{dq}{dt} \frac{1}{A}$

$$\frac{dq}{dx} = -kA \frac{\partial T}{\partial x}$$

## 1st Law of Thermodynamics / Energy

$$dE = dQ + dW$$

$$dU = TdS - PdV + \mu dn$$

( $\mu$  = chemical potential)

## The Partition Function

$$Z = \sum_j g_j e^{-\epsilon_j/kT}$$

$$S = \frac{U}{T} + Nk(\ln Z - \ln N + 1)$$

$$F = -NkT(\ln Z - \ln N + 1)$$

$$U = NkT^2 \left( \frac{\partial \ln Z}{\partial T} \right)$$

$$H = U + PV \quad dH = TdS - PdV$$

$$F = U - TS \quad dH = Tds + Vdp$$

## RMS Speed: quick derivation

$$\frac{3}{2}kT = \frac{1}{3}mv^2$$

$\sqrt{v_{rms}} \approx \sqrt{\frac{3kT}{m}}$

## Differentiation or Graham's Law of Effusion

$$\frac{\text{rate}_1}{\text{rate}_2} = \sqrt{\frac{M_2}{M_1}} \leftarrow \text{molar mass}$$

can be derived by simply noting that  $v_{\text{rms}} \propto \sqrt{\frac{T}{m}}$

# Quantum Physics (12/1)

## The Schrödinger Equation

$$\boxed{-\frac{\hbar^2}{2m} \nabla^2 \Psi + V\Psi = i\hbar \frac{\partial \Psi}{\partial t}}$$

### Operators

position  $\hat{x} = x$   
 momentum  $\hat{p} = -i\hbar \frac{\partial}{\partial x}$   
 Energy  $\hat{E} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2}$

### DeBroglie Formulae

$$E = \frac{p^2}{2m}$$

$$p = \sqrt{2mE}$$

$$p = \hbar k = \frac{\hbar}{\lambda}$$

### Planck Formulae

$$E = \hbar\omega = hf = \frac{hc}{\lambda}$$

only applies to photons!

### Heisenberg's Uncertainty Relations

$$\sigma_x \sigma_p \geq \frac{\hbar}{2}$$

$$\sigma_E \sigma_F \geq \frac{\hbar}{2}$$

### Ehrenfest's Theorem

$$\frac{d\langle p \rangle}{dt} = -\frac{\partial V}{\partial x}$$

### Compton Scattering Formula

$$\Delta\lambda = \frac{hc}{mc^2} (1 - \cos\theta)$$

$$\lambda_{\text{compton}} = \frac{hc}{mc^2} = \frac{h}{mc}$$

### TISE

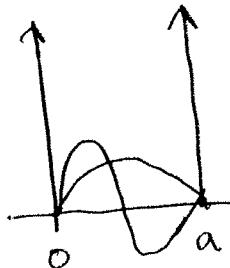
$$\boxed{-\frac{\hbar^2}{2m} \nabla^2 \psi + V\psi = E\psi}$$

### Time Dependence

$$f(t) = e^{-iEt/\hbar}$$

## Quantum Physics (Continued)

### Infinite Square Well



$$E = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2 \pi^2 n^2}{2ma^2}$$

$$\Psi = \sqrt{\frac{a}{\alpha}} \sin\left(\frac{n\pi}{a} x\right) e^{-iEt/\hbar}$$

~~Exercises~~

### The Harmonic Oscillator

$$\frac{-\hbar^2}{2m} \nabla^2 \Psi + V\Psi = E\Psi$$

$$\frac{-\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \Psi + \frac{1}{2} m\omega^2 x^2 \Psi = E\Psi \quad \omega^2 = \frac{k}{m}$$

$$\frac{1}{2m} \left[ \left( \frac{\hbar}{i} \frac{\partial}{\partial x} \right)^2 + (m\omega x)^2 \right] \Psi = E\Psi$$

$$a_{\pm} = \frac{1}{\sqrt{2m}} \left( \frac{\hbar}{i} \frac{\partial}{\partial x} \pm im\omega x \right)$$

$$(a_- a_+ - \frac{1}{2}\hbar\omega) \Psi = E\Psi$$

:

$$\Psi_0 = A_0 e^{-\frac{m\omega}{2\hbar} x^2}$$

④  $E_n = \left(n + \frac{1}{2}\right)\hbar\omega$

### The Free Particle

$$\Psi(x, t) = A e^{ik(x - \frac{\hbar k}{2m}t)} + B e^{-ik(x + \frac{\hbar k}{2m}t)}$$

Fourier transform:

$$\phi(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \Psi(x, 0) e^{-ikx} dx$$

### Group Velocity

$$V_{\text{group}} = \frac{\partial \omega}{\partial k}$$

### Phase Velocity

$$V_{\text{phase}} = \frac{\omega}{k}$$

### Delta Function

$$\int_{-\infty}^{\infty} f(x) \delta(x-a) dx = f(a)$$

### Basic Definitions

Hermitian operator:  $T^\dagger = T$

Unitary operator:  $U^\dagger = U^{-1}$

### Probability Current

$$\text{Prob. } \frac{dP_{ab}}{dt} = J(a, t) - J(b, t)$$

$$J(x, t) = \frac{ie\hbar}{2m} \left( \Psi \frac{\partial \Psi^*}{\partial t} - \Psi^* \frac{\partial \Psi}{\partial t} \right)$$

### Commutator Relations

$$[A, B] = [-B, A]$$

$$\begin{aligned} B, Ac &= BAc - ACB \\ A[Bc] &+ [B, A]c \\ ABC - ACB + BAc - ABC &= 0 \end{aligned}$$

$$A[Bx] - [Ax, B] = 2Ax - 2Ax = 0$$

$$[A, B] = AB - BA$$

$$[AB, C] = A[B, C] + [A, C]B$$

## Spin & Angular Momentum

$$\vec{L}^2 \psi = \hbar^2 l(l+1) \psi$$

$$L_z \psi = m\hbar \psi$$

## Pauli matrixes

$$\sigma_1 = \sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

$$\sigma_2 = \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

$$\sigma_3 = \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

## Hydrogen atom

$$E = -13.6 \left( \frac{1}{n^2} - \frac{1}{r^2} \right)$$

To derive, assume that  $mv r = n\hbar$

$$\text{Bohr radius } a_0 = \frac{4\pi\epsilon_0\hbar^2}{me^2}$$

## Angular Momentum quantum numbers

$n$ : principal quantum number -

$l$ : Azimuthal quantum number - gives orbital angular momentum

$m_l$ : magnetic quantum number, ranges from  $-l$  to  $l$

$s$ : spin quantum number

$$\text{for electrons } = \{1/2, -1/2\} \quad s_z = m_s \hbar$$

$$|\vec{S}| = \sqrt{s(s+1)} \hbar \quad \cancel{\text{for } s = m_s \hbar} = m_s \hbar$$

$= \frac{1}{2} \hbar \text{ for electrons.}$

$m_s$  goes from  
 $-s$  to  $s$  in  $\mathbb{Z}$

$l$	$\# e^-$	$2(2l+1)$
0	s	2
1	p	6
2	d	10
3	f	14

# Misc. Modern Physics Equations

## Positronium

$$\mu = \frac{m_e m_e}{m_e + m_e} = \frac{m_e^2}{2m_e} = \frac{m_e}{2} \Rightarrow E \text{ levels are } \frac{1}{2} \text{ of hydrogen's.}$$

## The Planck Length

$$L_p = \sqrt{\frac{G \hbar}{c^3}}$$

## Schwarzschild Radius

$$R_s = \frac{2MG}{c^2}$$

## Constants

$$kT \approx \frac{1}{40} \text{ eV at } 300\text{K}$$

$$hc = 1240 \text{ eV} \cdot \text{nm}$$

$$G = 6.67 \times 10^{-11}$$

$$K = 8.62 \times 10^{-5} \text{ eV/K}$$

## Photoelectric Effect

$$E = eV_{stop} = hf - \phi$$

## Bohr Atom

$$\Delta E = -Z^2 13.6 \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

## Spectroscopic Notation / "Term Symbol"

$$2s+1$$

$$L_J$$

$$s = \text{spin } \#$$

$$L = \ell = \begin{matrix} 0 & s \\ 1 & p \\ 2 & d \\ 3 & f \end{matrix}$$

$$L = \ell = \begin{matrix} s \\ p \\ d \\ f \end{matrix}$$

$\hat{J} = \text{total angular momentum}$

$$\hat{J} = \hat{s} + \hat{\ell}$$

$$L = \ell_1 + \ell_2 + \dots$$

$$s = s_1 + s_2 + \dots$$

for example  
 $3S$   
refers to  
 $2s+1=3$   
 $s=1$   
 $J=0+1=1$

## Misc.

### Intrinsic Magnetic Moment

$$\vec{\mu}_s = \gamma \vec{s}$$

$$\gamma = \text{gyromagnetic ratio} = \frac{e g}{2m}$$

### A useful rule of thumb for Energy levels

Electronic transitions  $\approx 1\text{ eV}$

Vibrational transitions  $\approx .1\text{ eV}$

Rotational transitions  $\approx .001\text{ eV}$

## Special Relativity (6/.)

$$\gamma = \frac{1}{(1 - \beta^2)^{1/2}}$$

$$\beta = \frac{v}{c}$$

$$\gamma(\beta = 1/2) = 1.15$$

$$\gamma(\beta = .9) = 2.3$$

## Energy-momentum

$$E^2 = p^2 c^2 + m^2 c^4$$

\*master equation\*

$$E = \gamma mc^2$$

$$\frac{mc^2}{(1-\beta)}$$

$$KE = \gamma mc^2 - mc^2$$

$$p = \gamma mv$$

for a photon,  $m=0$  and

$$E = pc$$

## Lorentz transformation

$$\begin{aligned}x' &= \gamma(x - vt) \\y' &= y \\z' &= z \\t' &= \gamma\left(t - \frac{xv}{c^2}\right)\end{aligned}$$

## \* Transformation of velocities \*

$$\underline{u} = \frac{dx}{dt} = \text{velocity in frame } S$$

$$u' = \frac{dx'}{dt'} = \text{velocity in frame } S'$$

$S'$  is moving with velocity  $v$

$$u' = \frac{u - v}{1 - \frac{uv}{c^2}}$$

## \* The invariant / spacetime interval \* proper distance \*

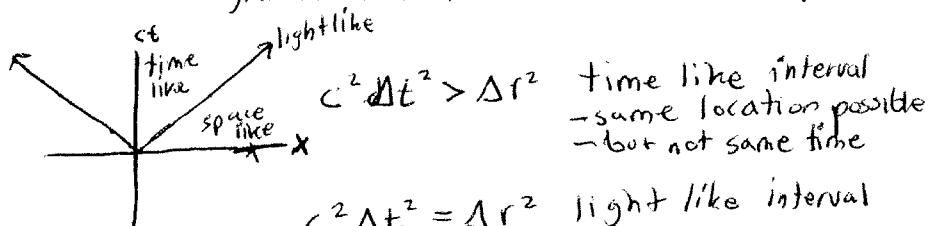
$$ds^2 = dr^2 - c^2 dt^2$$

$$4\text{vectors: } \underline{R}_a = \begin{pmatrix} ct \\ \underline{x} \end{pmatrix} = \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix}$$

$$\text{dot product } \underline{v}_1 \cdot \underline{v}_2 = ct_1 ct_2 - \underline{x}_1 \cdot \underline{x}_2$$

$$\text{Energy-momentum 4 vector } \underline{p} = \begin{pmatrix} E \\ p_x c \\ p_y c \\ p_z c \end{pmatrix}$$

The length of 4-vectors is invariant.



$$c^2 dt^2 = dr^2 \text{ light like interval}$$

$$c^2 dt^2 < dr^2 \text{ space-like interval}$$

- same time possible
- but not same location

## \* Length contraction \*

$$l = \frac{l_0}{\gamma}$$

## The Doppler Shift

$$E' = \gamma E_0 + \beta \gamma E =$$

Blue shift:

$$\frac{\lambda'}{\lambda} = \sqrt{\frac{1-\beta}{1+\beta}}$$

Red shift:

$$\frac{\lambda'}{\lambda} = \sqrt{\frac{1+\beta}{1-\beta}}$$

# Laboratory

## Standard deviation

$$\sigma^2 = \overline{(x - \bar{x})^2} = \langle (x - \langle x \rangle)^2 \rangle = \boxed{\langle x^2 \rangle - \langle x \rangle^2}$$

Poisson distribution (random distribution)

$$\sigma \approx \sqrt{\bar{x}}$$

## Propagation of errors

$$* \quad \boxed{\delta_x = [\delta_{x_1}^2 + \delta_{x_2}^2 + \dots]^{1/2}} *$$

Random errors : can be seen during multiple measurements and  
thus accounted for

Systematic errors : errors that are intrinsic to the instrumentation  
and cannot be revealed by repeated measurement.

# Particle Physics (3/.)

Hadrons : composite particles made of quarks

↳ Mesons : 2 quarks

~~B~~  $\pi^+, \pi^0, K^+, K^0, \psi$

↳ Baryons : 3 quarks

$p = udd$

$n = udd$

$\Lambda^0, \Sigma^+$

Baryon  $\text{#}$  = +1 for baryons,  
-1 for anti-baryons  
0 for non-baryons

Fermions : 1/2 integer spin

↳ Leptons : spin 1/2

$e^-, \mu^-, \tau^-, \nu_e, \nu_\mu, \nu_\tau$

There are 3 lepton  $\text{#}$ s

$L_e, L_\mu, L_\tau = \{+1, -1, 0\}$

↳ Quarks : spin 1/2

$+2/3 U \quad +2/3 C \quad +2/3 T$

$-1/3 D \quad -1/3 S \quad -1/3 B$

Bosons : integer spin particles

$\gamma, g, \text{graviton}, W^+, W^-, Z^0$

Conservation Laws :

Energy & momentum

Parity - not in weak

charge

Time symmetry

Strangeness - not in weak

Baryon  $\text{#}$ .

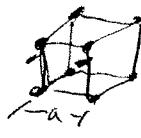
Lepton  $\text{#}$

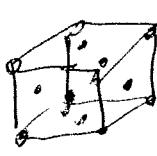
# Solid State Physics ( $\sim 3/1$ )

## The Primitive Cell

$$V_{\text{primitive cell}} = \frac{V_{\text{Bravais lattice}}}{\# \text{ of lattice points}}$$

 simple cubic  $\rightarrow V_p = a^3$   
1 lattice point inside

 B.C.C.  $\rightarrow V_p = a^3/2$   
2 lattice points

 F.C.C.  $\rightarrow V_p = a^3/4$   
4 lattice points

## Resistivity of a semiconductor

$$\rho \approx 1/T$$

## Effective Mass

$$m^* = \frac{\hbar^2}{\frac{\partial^2 E}{\partial k^2}}$$

"really simple" derivation:

$$\begin{aligned} p &= \hbar k \\ E &\approx \frac{p^2}{2m} = \frac{\hbar^2 k^2}{2m} \\ \frac{\partial E}{\partial k^2} &= \frac{\hbar^2}{m} \quad \frac{1}{m} = \frac{\frac{\partial E}{\partial k^2}}{\hbar^2} \\ m &= \frac{\hbar^2}{\frac{\partial^2 E}{\partial k^2}} \end{aligned}$$

## Math ( $\sim 3'$ )

mapping

gradient  $\nabla f = \left( \frac{\partial f_1}{\partial x_1}, \dots, \frac{\partial f_n}{\partial x_n} \right) \quad \mathbb{R}^n \rightarrow \mathbb{V}^n$

divergence  $\vec{\nabla} \cdot \vec{F} = \frac{\partial F_1}{\partial x_1} + \dots + \frac{\partial F_n}{\partial x_n} \quad \mathbb{V}^n \rightarrow \mathbb{R}$

Curl (in  $\mathbb{V}^3$ )  $\vec{\nabla} \times \vec{F} = \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ F_x & F_y & F_z \end{vmatrix} \quad \mathbb{V}^n \rightarrow \mathbb{V}^n$

## Divergence Theorem

$$\int_{\partial V} \vec{F} \cdot d\vec{S} = \int_V \vec{\nabla} \cdot \vec{F} dV$$

## Stoke's Theorem (basic form for physics)

$$\int_{\partial A} \vec{F} \cdot d\vec{r} = \int_A \vec{\nabla} \times \vec{F} \cdot d\vec{A}$$

## Important Identities

$$\nabla \cdot (\nabla \times \vec{H}) = 0$$

$$\nabla \times (\nabla f) = 0$$

## BAC - CAB rule

$$A \times (B \times C) = B(A \cdot C) - C(A \cdot B)$$

# Basic Equations from Astronomy / Astrophysics ( $\sim 1/1$ )

$$m_1 - m_2 = -2.5 \log_{10}(F_1/F_2)$$

Apparent magnitude

$$M = m + 5 - 5 \log_{10}(d/\text{pc}) - A$$

Absolute magnitude

$\begin{matrix} T_{\text{stellar}} \\ (\text{extinction}) \end{matrix}$

$$f_{\text{eq}} \propto \frac{1}{f}$$

$$\text{Magnification} = f_o/f_{\text{eyepiece}}$$

$$\alpha_c = \frac{1.22\lambda}{D}$$

Angular resolution / Rayleigh criterion

$$\text{Image scale} = \frac{1}{f_o \tan i''} = \frac{206,265}{f} \Rightarrow \frac{\text{arcseconds}}{\text{mm}}$$

## Basic tips

- If a problem looks hard / tedious, try process of elimination first.
- Use limiting cases to eliminate answers.

- Very few problems are simply "plug into this equation" although a few are. Most problems combine at least 2 concepts or have a twist ~~or~~ to them.
- Dimensional analysis is occasionally useful.

### Statistics from past Exams

Exam	# needed to get 800	% got 800	% got 800	% got 800	# to get 900	% got 900	% got 900	% got 900
GRE 8677 1981-1984	45	33%	59	15%	72	5%	84	2%
GRE 9677 1993-1996	32	39%	44	21%	56	9%	67	3%
GRE 9277 1988-1991	40	42%	53	23%	65	10%	76	3%
GRE 0177 2000-2003	44	41%	58	22%	73	10%	85	2%
Average	41	39%	53	20%	67	8.5%	78	2.5%

more difficult questions

more difficult questions